

Use of a Night-Tracking during Daytime

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AIUB

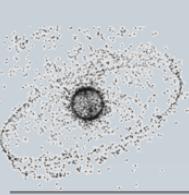
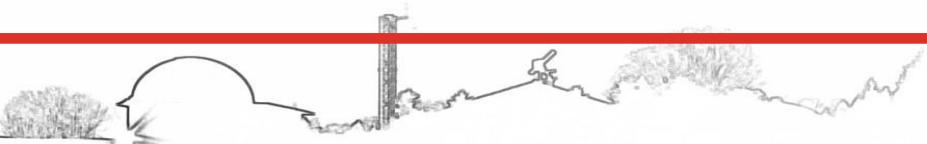


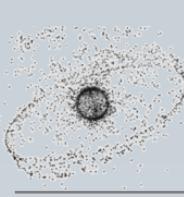
Table of Contents

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- Motivation
- Needed Hardware
- What we should expect?
 - And what can we improve?
- Results
 - Observation results
 - Orbit Determination
 - Light Curves
- Conclusion & Outlook





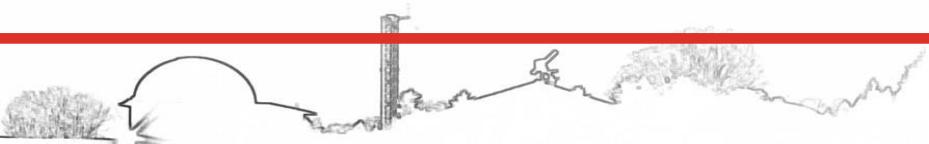
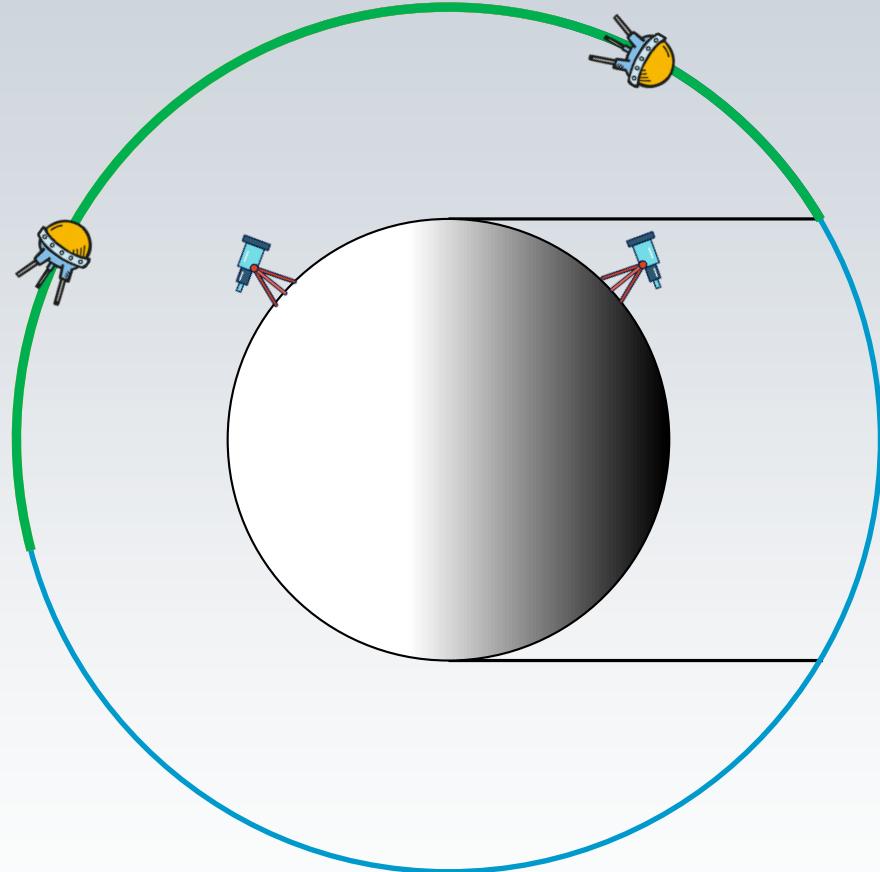
1st Motivation: Orbit Determination

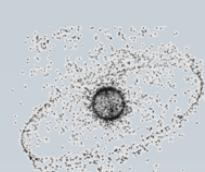
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Space Debris Tracking Case:

- 1 Single observatory
 - Only SLR
 - Only Optical
 - Both (even better)
- Usual case:
 - Observer in Earth shadow
 - Illuminated target
 - Observable Portion of Orbit
- Daytime observation
 - Observer and target in Sunlight
 - Observable Portion of Orbit
- Extended observation scenario (both night- and daytime)
 - Increased observable portion of orbit!!!
 - ⇒ Increase of orbit determination accuracy!!!

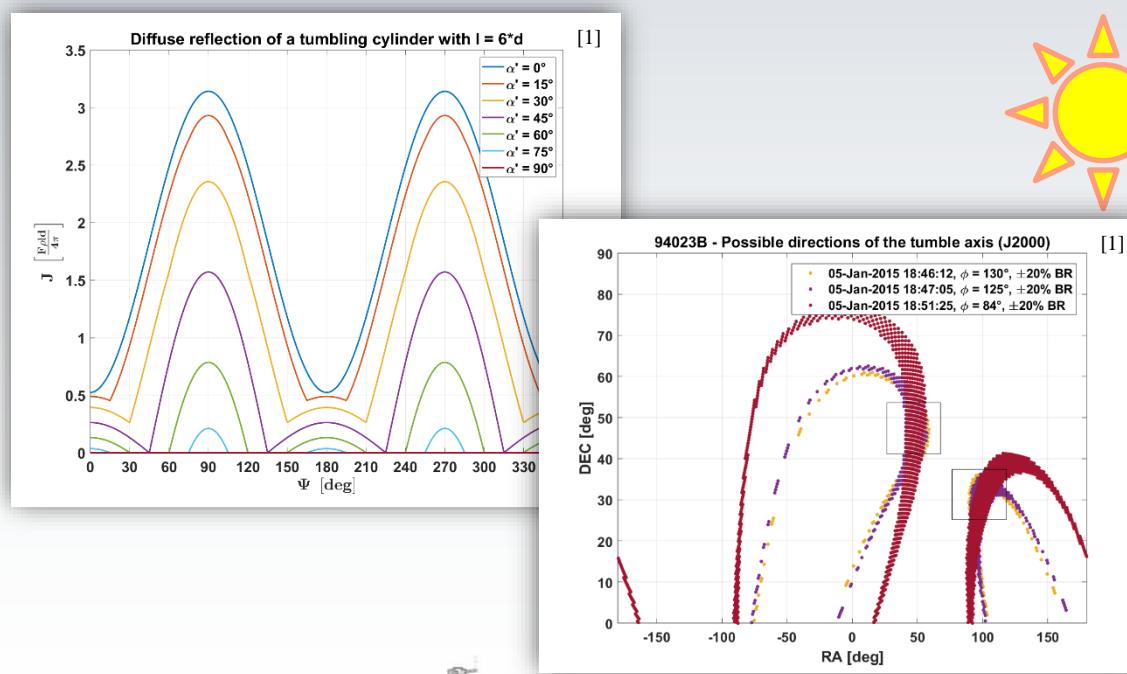


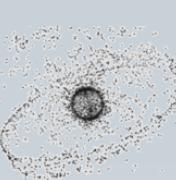
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2nd Motivation: Attitude Determination

Space Debris Attitude determination from light curves:

- Object Spin: stable into the inertial system
- Light curve features depend on illumination condition (α')
- Successful attitude determination
 - Enough observation geometry variation
 - Short time difference between observation series (for MEO and GEO, ok! And for LEO?)





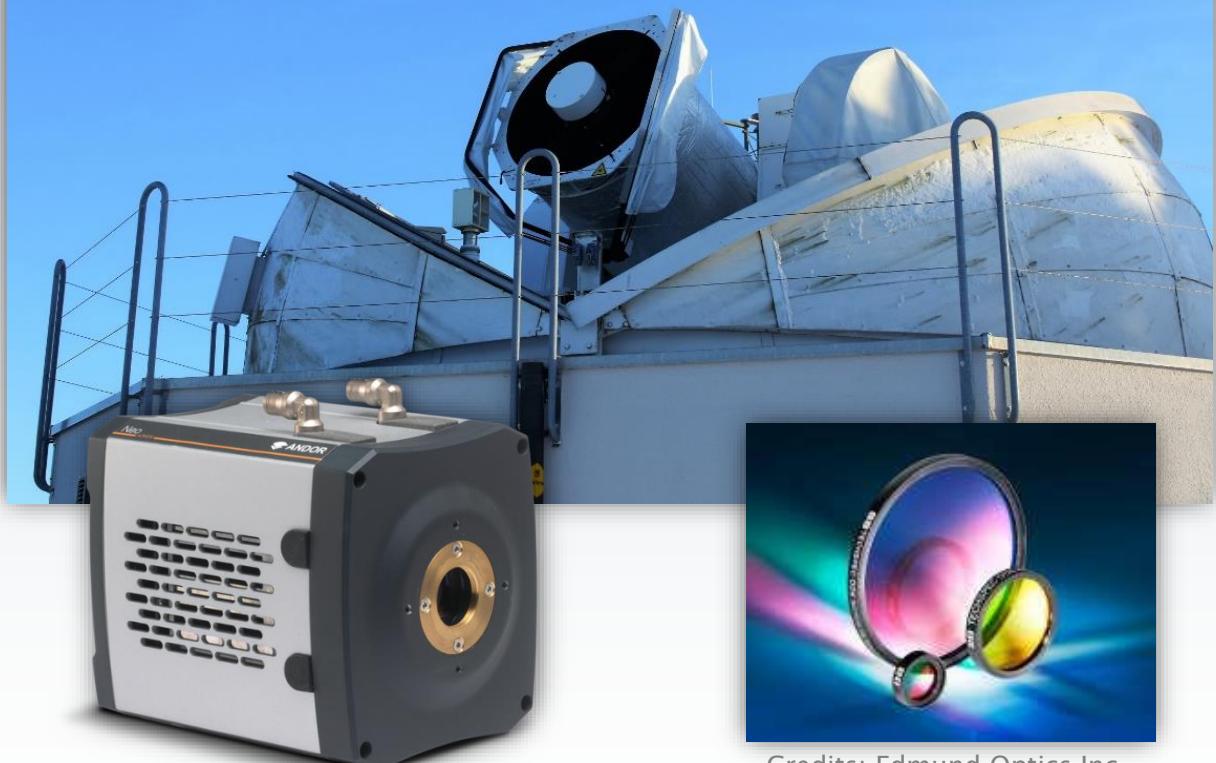
What do we use...

Tracking Camera Hardware

- **Telescope: ZIMLAT**
 - 1-m Aperture Ritchey-Crétien
 - Coudé focus for Laser
 - Nasmyth Focus for tracking and CCD cameras
(available focal lengths: 1.2m, 2x4m, **8m**)
- **Camera: Neo 5.5 sCMOS**
 - 1 e- read noise
 - TE cooling to -40°C
 - 5.5 megapixel sensor, $6.5 \mu\text{m}$ pixels
 - 22 mm diagonal field of view
 - Rolling and Global Shutter
 - Rapid frame rates
 - 30 fps over extended kinetic series
 - Burst to memory at 100 fps full frame
- **Filters**
 - **532nm, 25mm Dia., OD 6 Blocking Notch Filter (SLR)**
 - **500nm, 25mm Dia., OD 4 Longpass Filter (Sky)**

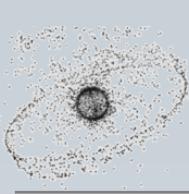
ZIMLAT

Zimmerwald Laser and Astrometry Telescope



Credits: Oxford Instruments 2018

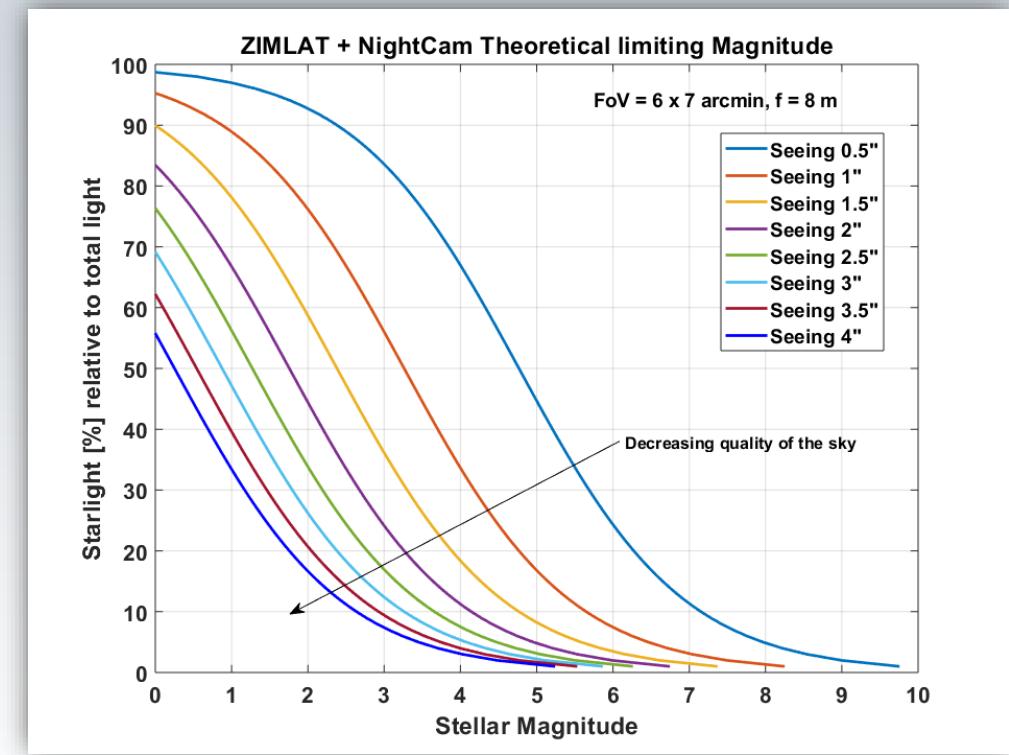
Credits: Edmund Optics Inc.



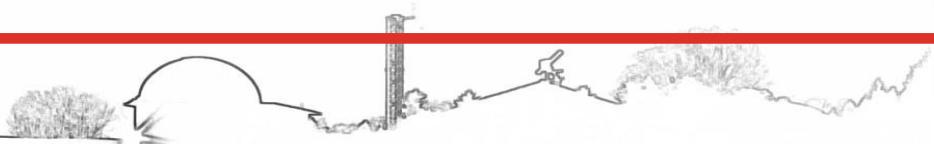
Daytime Limiting Magnitude

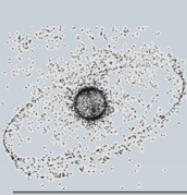
Factors influencing the limiting magnitude

- Sky background brightness (~3 Mag [1])
 - Distance from Sun
 - Distance from Zenith
- Astronomical Seeing
(Zimmerwald <2 arcsec night-time)
 - Humidity
 - Extinction
 - Temperature
 - Turbulences
- Focal length = 8 m, sensor diagonal = 22 mm
 - Pixel scale = 0.17 arcsec (0.0289 arcsec^2)
 - FoV = 6 x 7 arcmin
 - Affect the Signal to Noise Ratio of the observed object



[1] Steindorfer M., Kirchner G., Koidl F., Wang P., Kucharski D.. Recent Space Debris Related Activities at the SLR Station Graz. 1st NEO and Space Debris conference, 22-24 January 2019, Darmstadt, Germany.





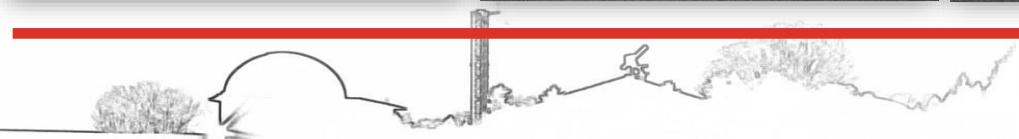
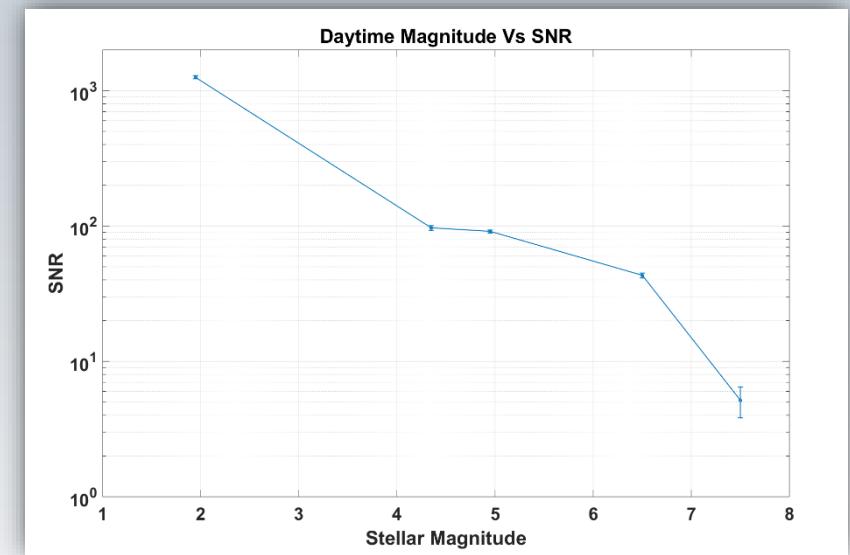
First Observations – Daytime Limiting Magnitude

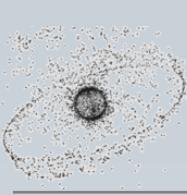
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Stars daytime observations

- Observations possible up to Magnitude 7
- Strong brightness gradient in the sky due to:
 - Distance from the Sun
 - Elevation
 - Weather conditions





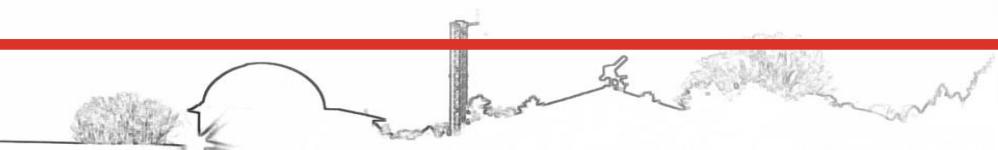
First Results – Satellite Observations

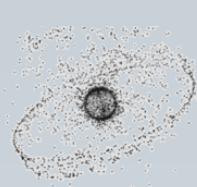
Daytime satellite tracking

- Telescope: ZIMLAT
- Target: PAZ
- Altitude: 507–510 km
- Exposure time: 0.1 seconds

Interesting outcomes

- Speckles and seeing effects on the object
- Increase of sky brightness over 1.5 minutes
- Influence of object elevation and distance from the observer

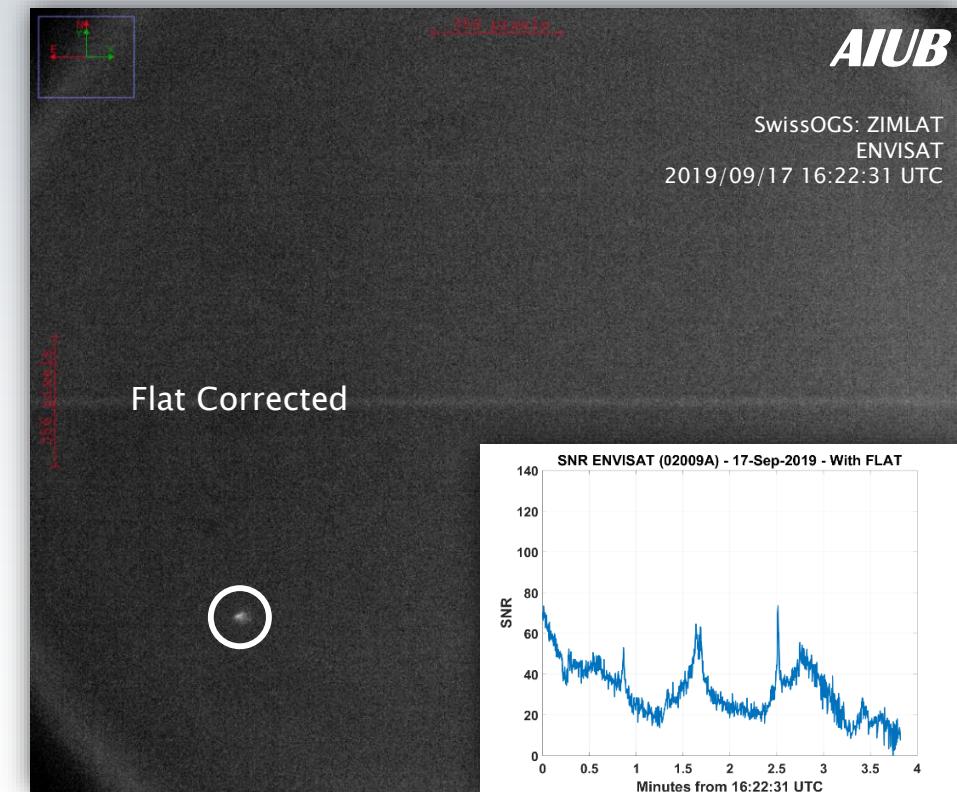
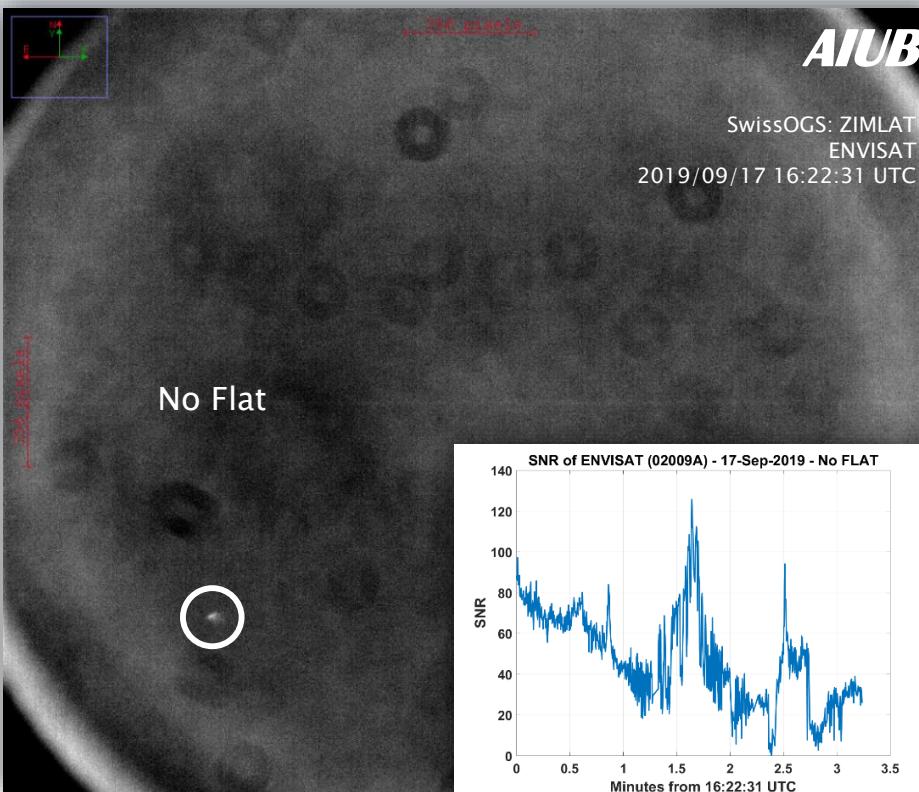




Improving Performances

Flat field Correction:

- Reduces SNR
- Reduces background noise
- Improves object detectability



Improving Performances

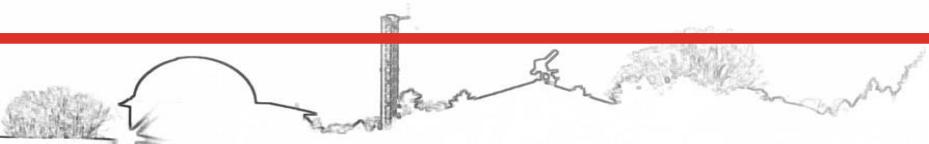
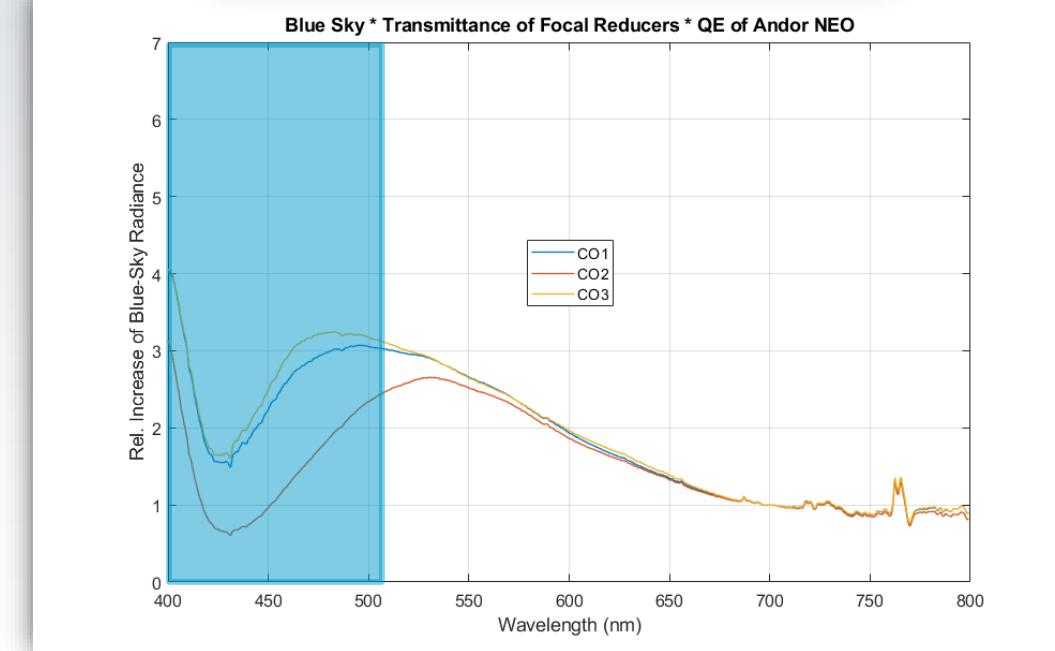
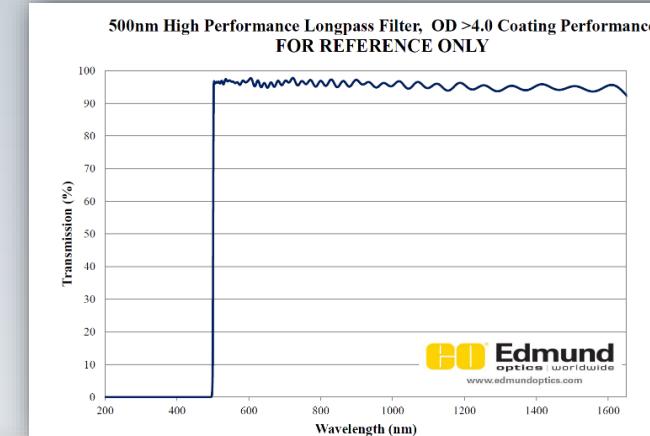
Removing daytime sky contribution

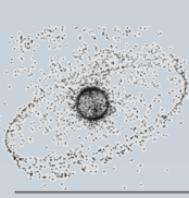
- Measured the sky spectrum
- Accounting for transmittancy of focal reducers
- Accounting of Camera Quantum Efficiency
- Longpass edge filter selection (Sky Filter)

	Without Sky filter	With Sky filter
Exp time	0.05	0.05
Star Magnitude	1.95	1.95
SNR	803.05 ± 47.42	818.20 ± 19.46
Sky (ADU)	34213.16 ± 6.52	21045.41 ± 4.35

Improvement of Sky filter (500 nm)

- ⇒ +1.88% SNR
- ⇒ -38.5% of Sky background (same exp. Time)
- ⇒ ~ Sky background value (96.57%), +60% Exp. time

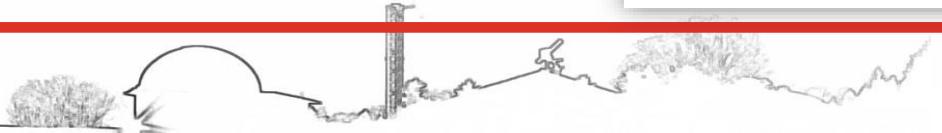
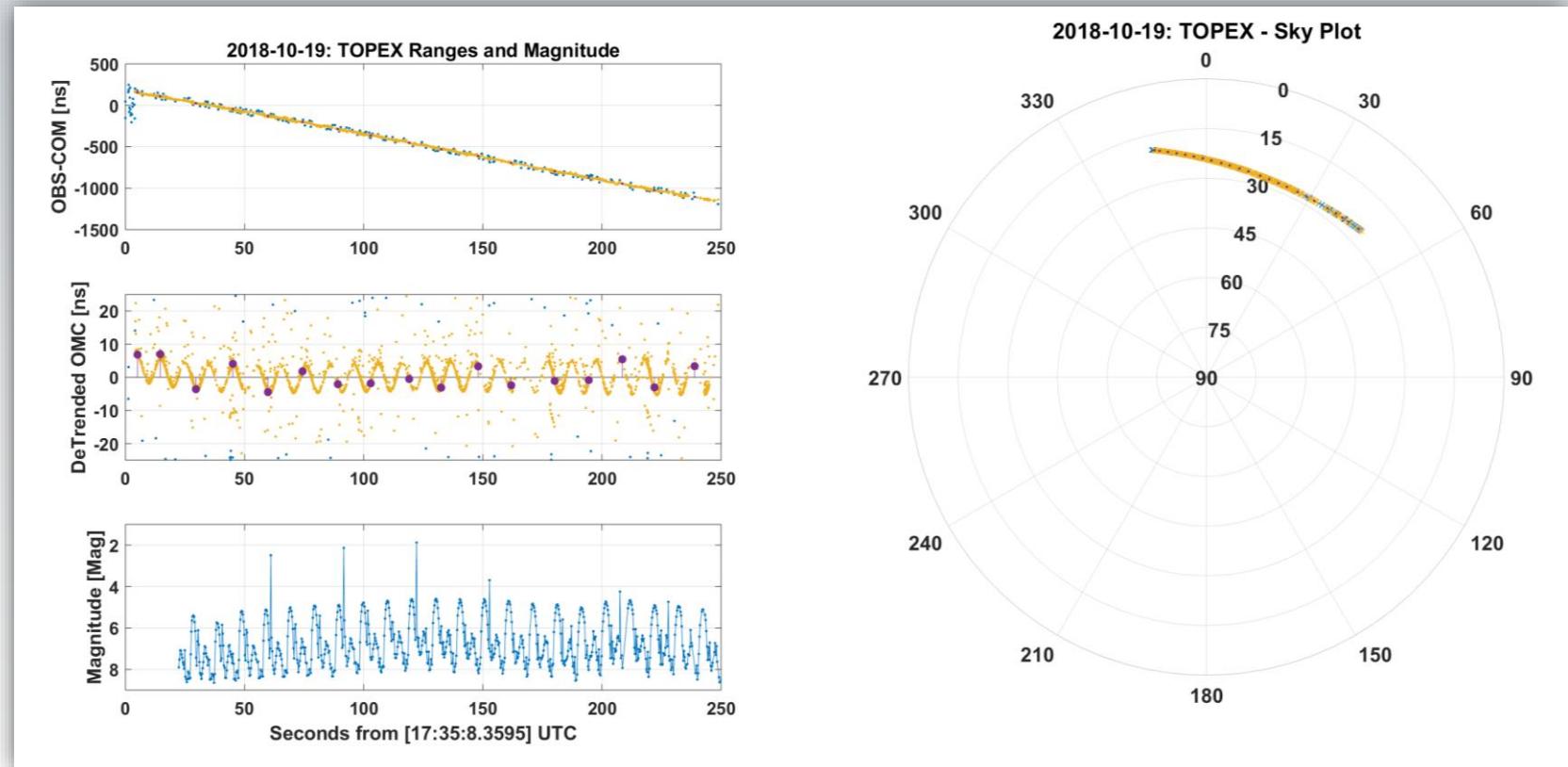




Tracking Camera Output Data

Night-Tracking Camera output:

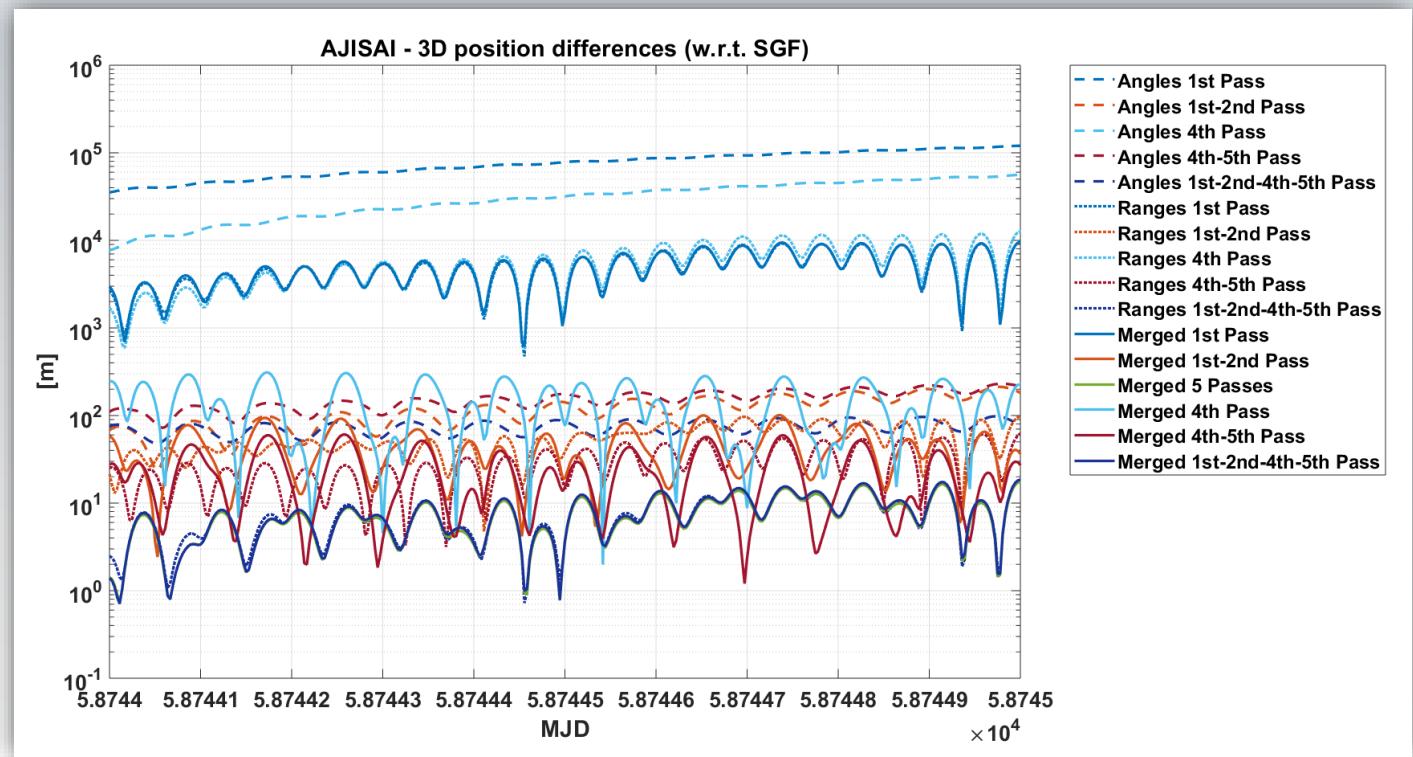
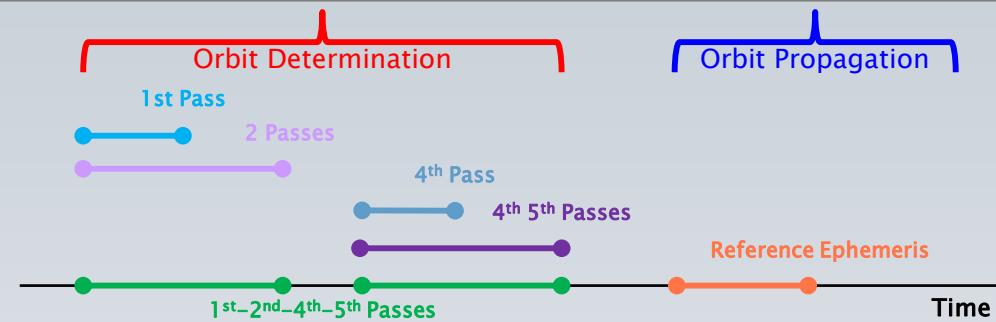
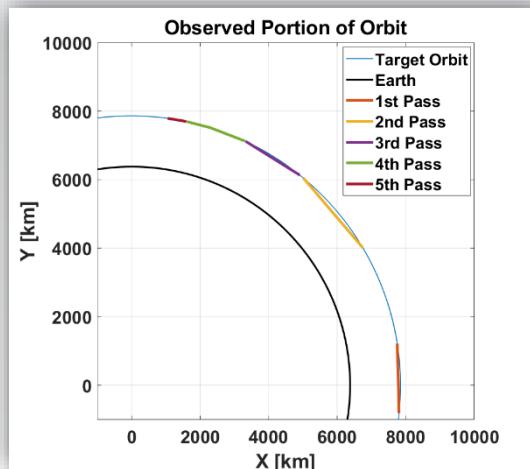
- Orbit determination
 - Angular direction of telescope pointing (encoders)
 - Laser Ranges
- Attitude Determination
 - Light curve
 - «Laser light curve»

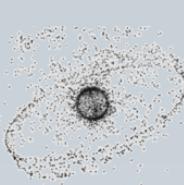


Orbit Determination Using Daytime Data

Ajisai Analysis

- Telescope: ZIMLAT
- Only Real Measurements
 - Angles from telescope encoder
 - Ranges from SLR
 - Total of 5 consecutive passes
- Date: 2019/09/17
 - 1st Pass 13:30 UTC
 - 2nd Pass 15:30 UTC
 - 3rd Pass 17:30 UTC
 - 4th Pass 19:30 UTC
 - 5th Pass 21:30 UTC





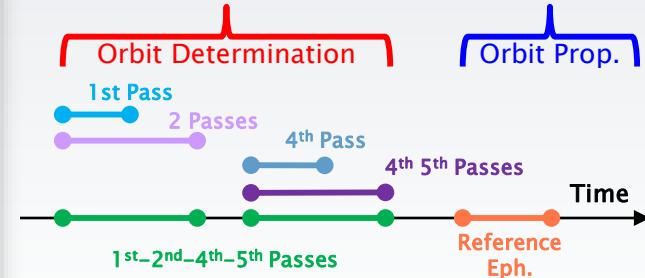
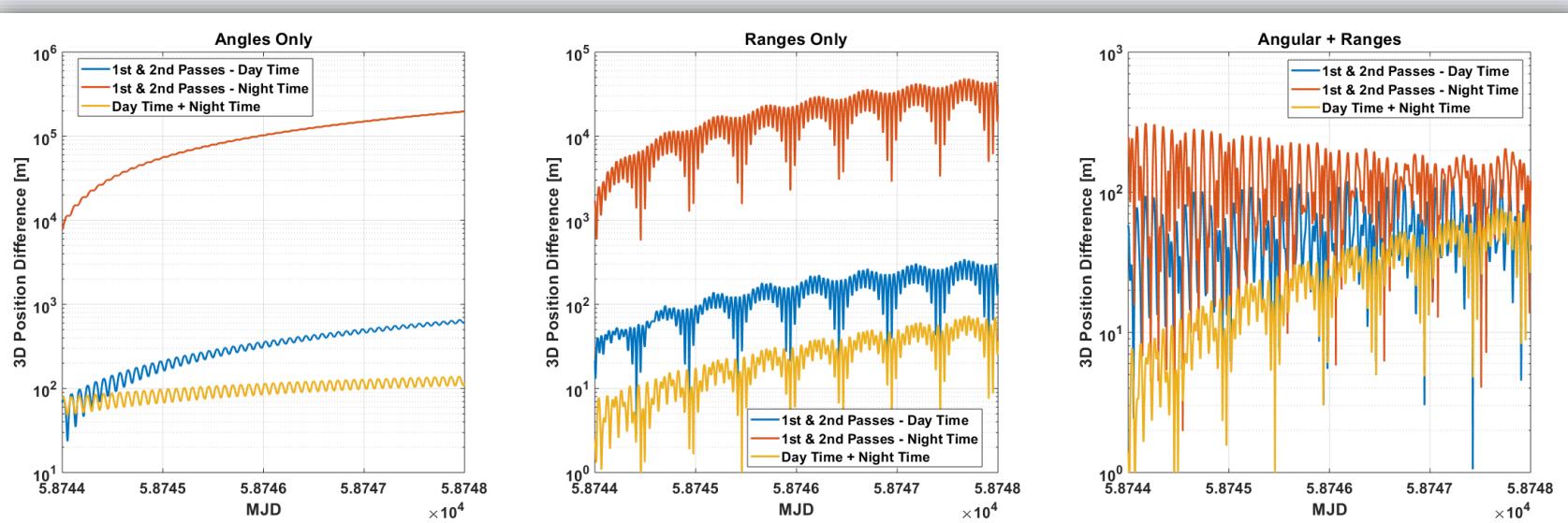
Orbit Determination Using Daytime Data

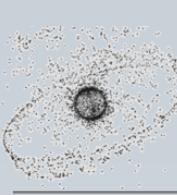
Ajisai Analysis – Playing with observables

Same observation scenario, comparison of accuracies obtainable by processing different observables

- Improvement of up to 2 order of magnitude given by increased portion of observation arc

Ajisai – 3D Position Difference w.r.t. SGF [m]			
Length of Observation arc	Angles Only	Ranges Only	Merged
1 Pass Daytime	202077.98	12938.65	12755.87
2 Passes Daytime	336.38	138.58	51.43
1 Pass Nighttime	102461.79	18393.05	126.34
2 Passes Nighttime	320.28	89.87	19.47
2 Daytime + 2 Nighttime	97.97	25.51	26.82
All available (5 passes)	115.43	25.34	25.49

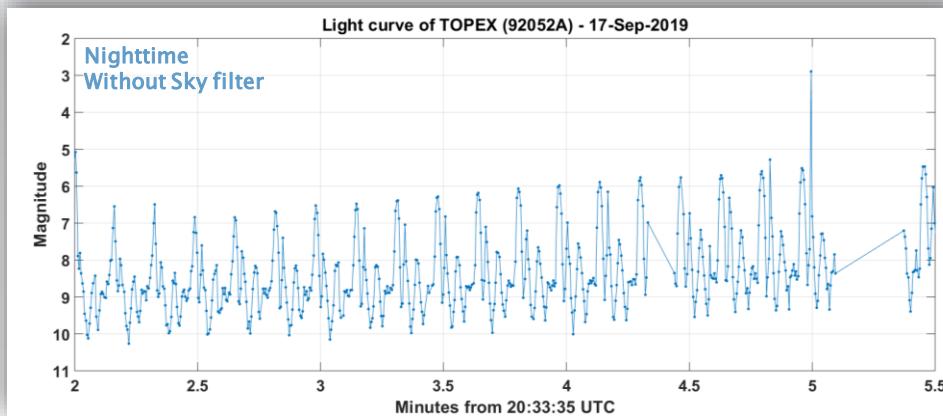
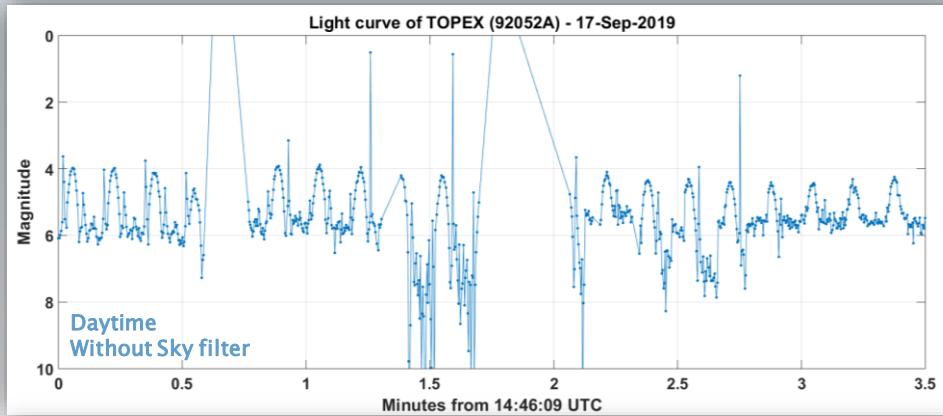




Observation for Attitude Determination

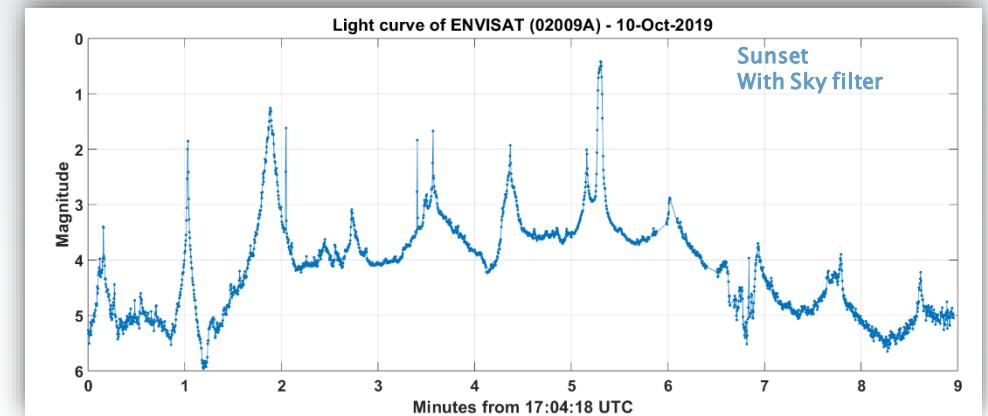
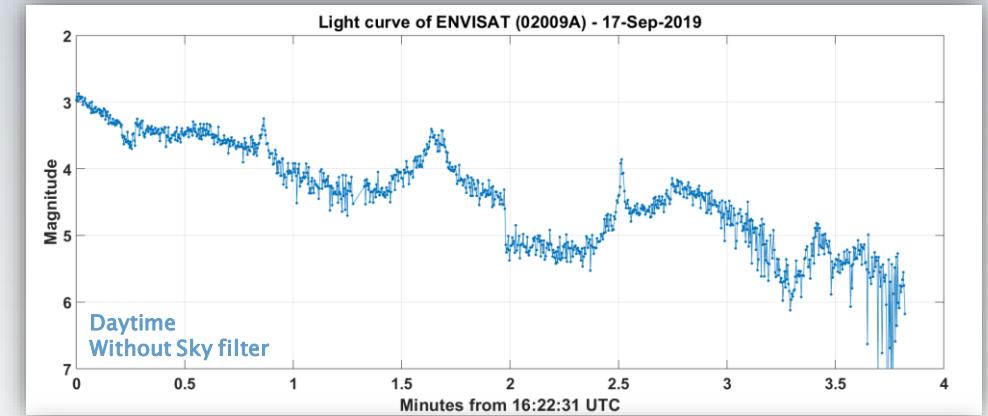
TOPEX

- Telescope: ZIMLAT
- Observation time:
2019/09/17 14:46 & 16:40 & 18:30 & 20:33 UTC
- Exp. Time: 0.05 seconds & 0.1 seconds
- Without Sky filter



ENVISAT

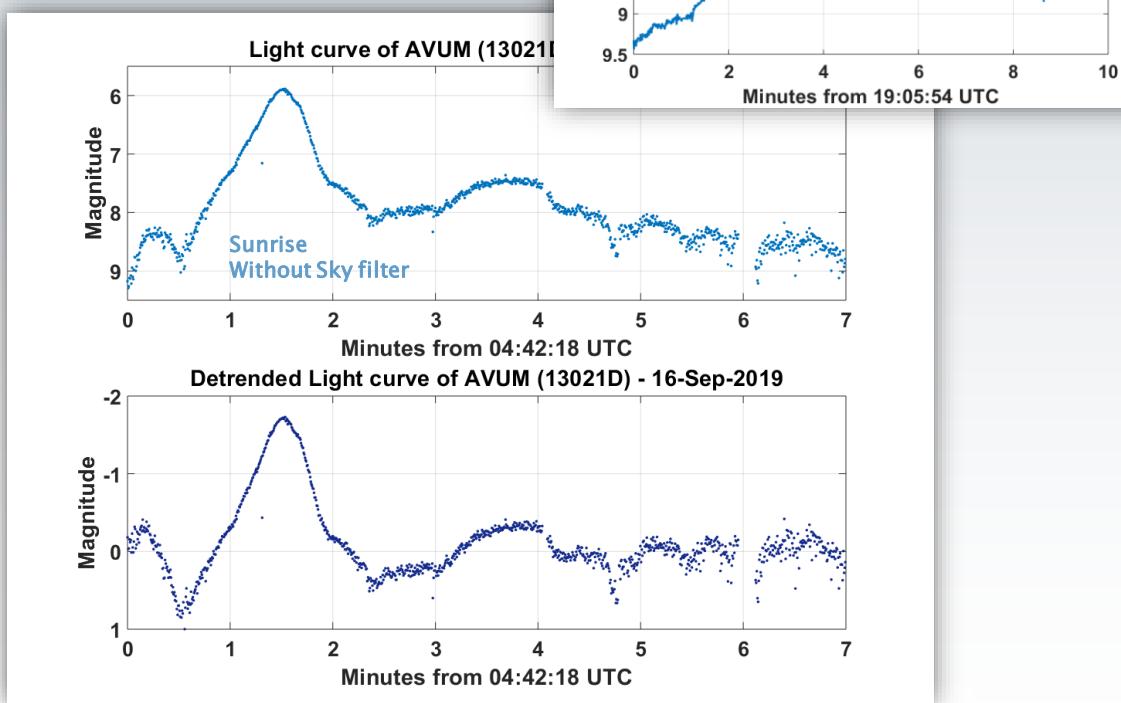
- Telescope: ZIMLAT
- Observation time:
 - 2019/09/17 16:25 UTC Without Sky filter
 - 2019/10/10 17:04 UTC With Sky filter
- Exp. Time: 0.01 seconds 0.05 seconds



Observation for Attitude Determination

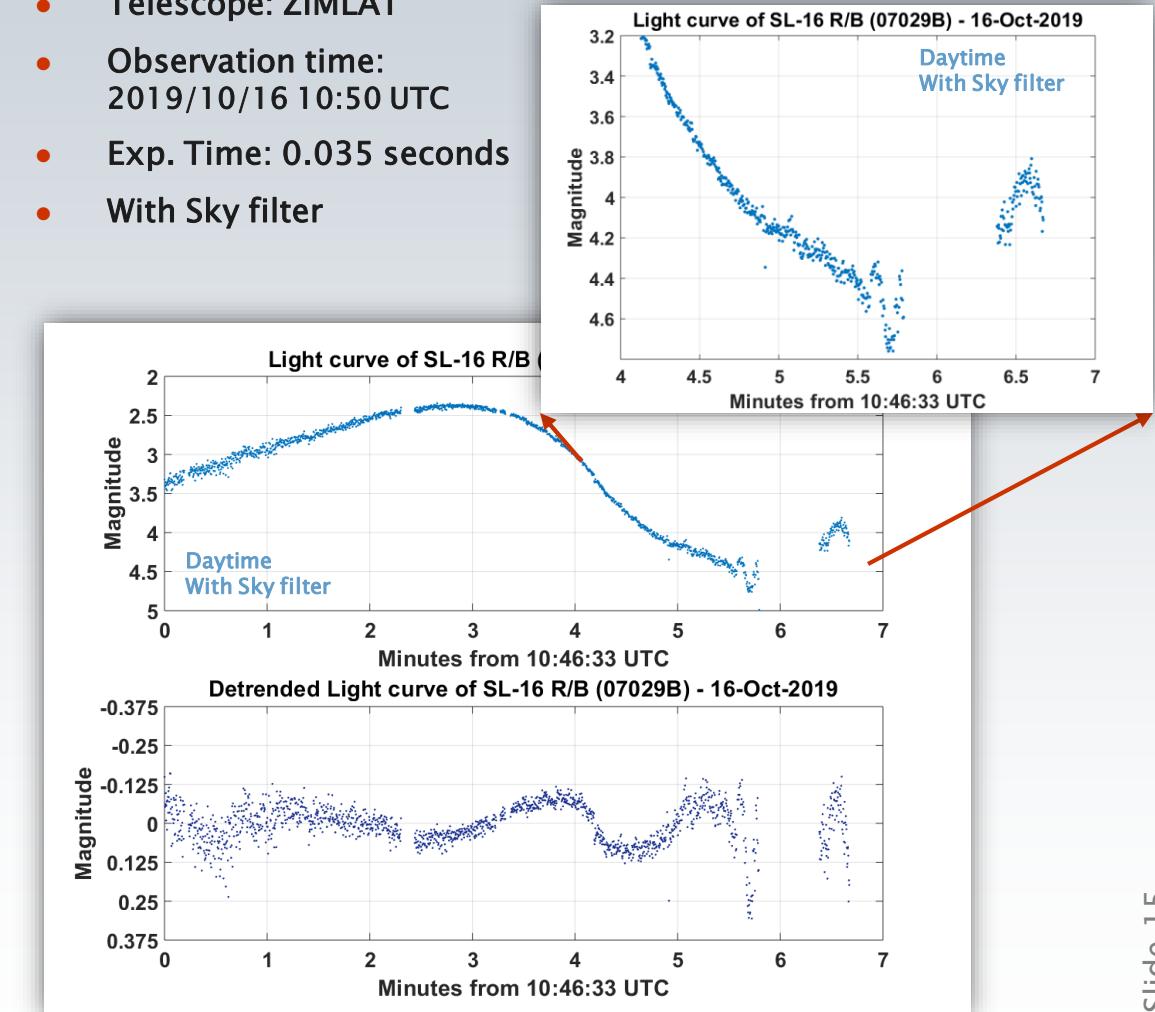
AVUM DEB (ADAPTOR)

- Telescope: ZIMLAT
- Observation time:
2019/09/15 19:10 UTC
2019/09/16 04:45 UTC
- Exp. Time: 0.2 seconds
- Without Sky filter



SL-16 R/B

- Telescope: ZIMLAT
- Observation time:
2019/10/16 10:50 UTC
- Exp. Time: 0.035 seconds
- With Sky filter

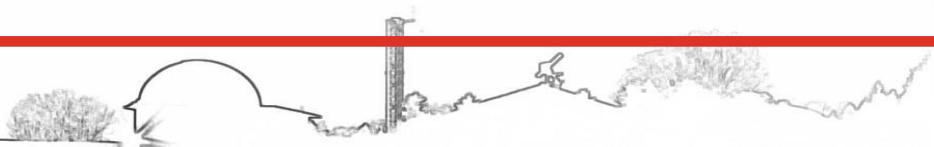


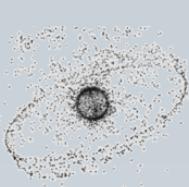
Summary

- Employment of the Night-Tracking Camera for Daytime observations
- Investigation of the limits of the system
 - Observation possible up to Magnitude 7 during daytime
- Evaluation of possible system improvement
 - Sky filter
 - Flat field correction
- Acquisition and analysis of measurements for:
 - Orbit determination
(up to 2 orders of magnitude of accuracy improvement)
 - Attitude determination
 - Active tracking
 - debris tracking
 - re-entering objects

Observed Objects during daytime		
Name	COSPAR ID	Altitude [km]
ENVISAT	02009A	800
TOPEX	92052A	1350
AJISAI	86061A	1485
AVUM DEB	13021D	664–800
SL-16 R/B	07029B	843–846
H-2A R/B	12025E	579–655
CZ-4C R/B	19066B	435–612
CZ-4B R/B	02024C	809–882
GEO-IK-2	16034A	943
HY-2A	11043A	971
HY-2B	18081A	971
CRYOSAT2	10013A	720
PAZ	18020A	507–510
TanDEM-X*	07026A	514

* Together with TerraSAR-X





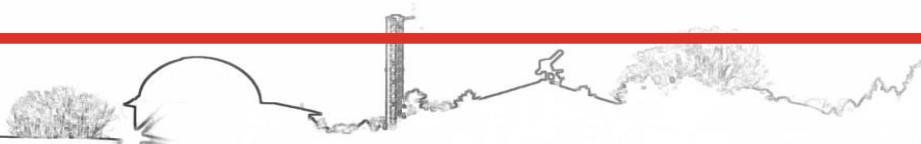
Summary & Outlook

Pros & Cons

- **Advantages of daytime observations:**
 - Increase of the length of the observation arc
 - Improve the achievable orbit determination accuracy
 - Big change in the illumination conditions
- **Disadvantages of daytime observations**
 - Limit to lower and bright objects
 - Diffuse sky-light degrade the image quality highlighting dust and defect on optical path
 - Strong background brightness changes due to
 - Weather
 - Distance from the Sun
 - Elevation of the object

Next Steps

- **Automation**
 - Test Object Recognition algorithm day- and nighttime
 - Adjustment of Image contrast
 - Ephemeris Correction
(from Az, El to Along-, Cross-track)
 - Object Planning
- **Real time Flat field?**
 - Could improve for object recognition
 - Improve post process analysis





Thank you for your attention